

Phase calibration in SMA

Daniel Espada Fernández

Harvard-Smithsonian Center for Astrophysics

Instituto de Astrofísica de Andalucía - CSIC

Marie Curie fellowship

- ASIAA (alphabetical order)
 - Derek Kubo, Sheng-Yuan Liu, Satoki Matsushita, Yu-Nung Su
- CfA:
 - Rob Christensen, Brian Koge, (T. Hunter), T.K. Sridharan, Paul Yamaguchi
- IAA:
 - Vicent Martinez, AMIGA group



Why gain calibration in mm/submm?

- ✦ Water vapor content fluctuations in the troposphere:
 - Absorption
 - Excess path delay (decorrelation of the phase)

SMA, pathfinder of ALMA

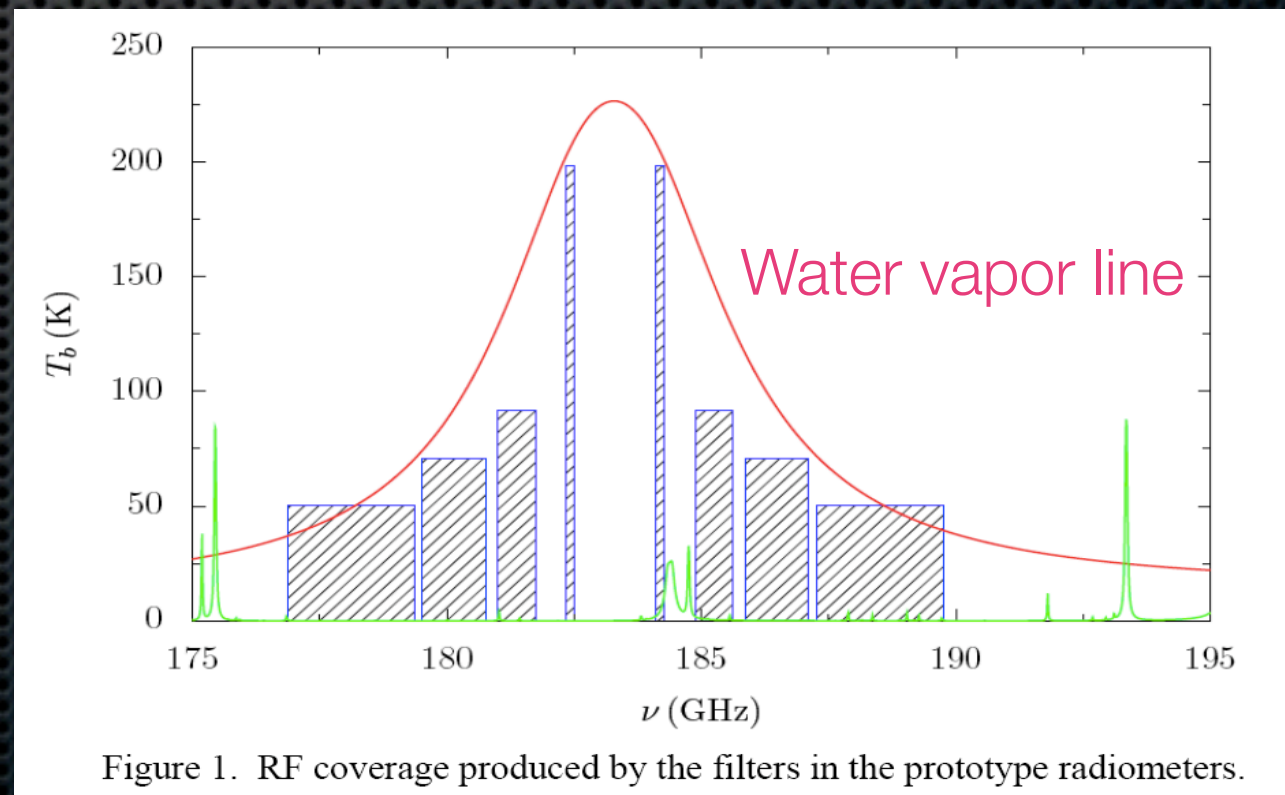
- ❖ SMA, 8 6m antennas. Mauna Kea (4000 m)
- ❖ 200,300,400 and 600 GHz Rxs.
- ❖ Dual-frequency mode
- ❖ 500 m baseline

Phase calibration techniques

- ✦ Water Vapor Radiometer
- ✦ Fast switching
- ✦ Phase Transfer

Phase calibration techniques

- ✦ **Water Vapor Radiometer** (e.g. #210, # 252, #496)
 - ✦ Sufficient system stability.
 - ✦ Good calibration to convert from water vapor to path delay.
 - ✦ Rely on atmospheric models.



Phase calibration techniques

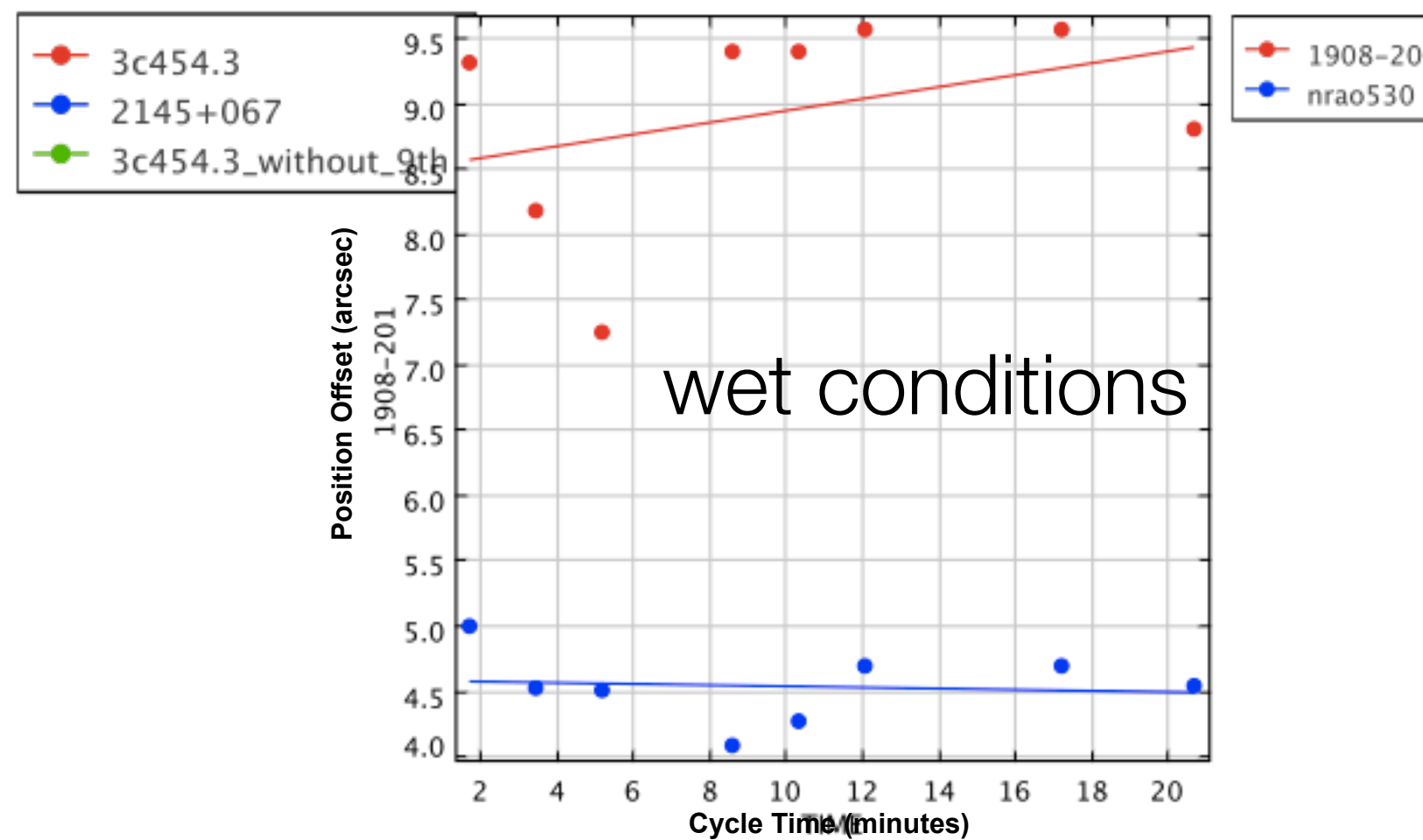
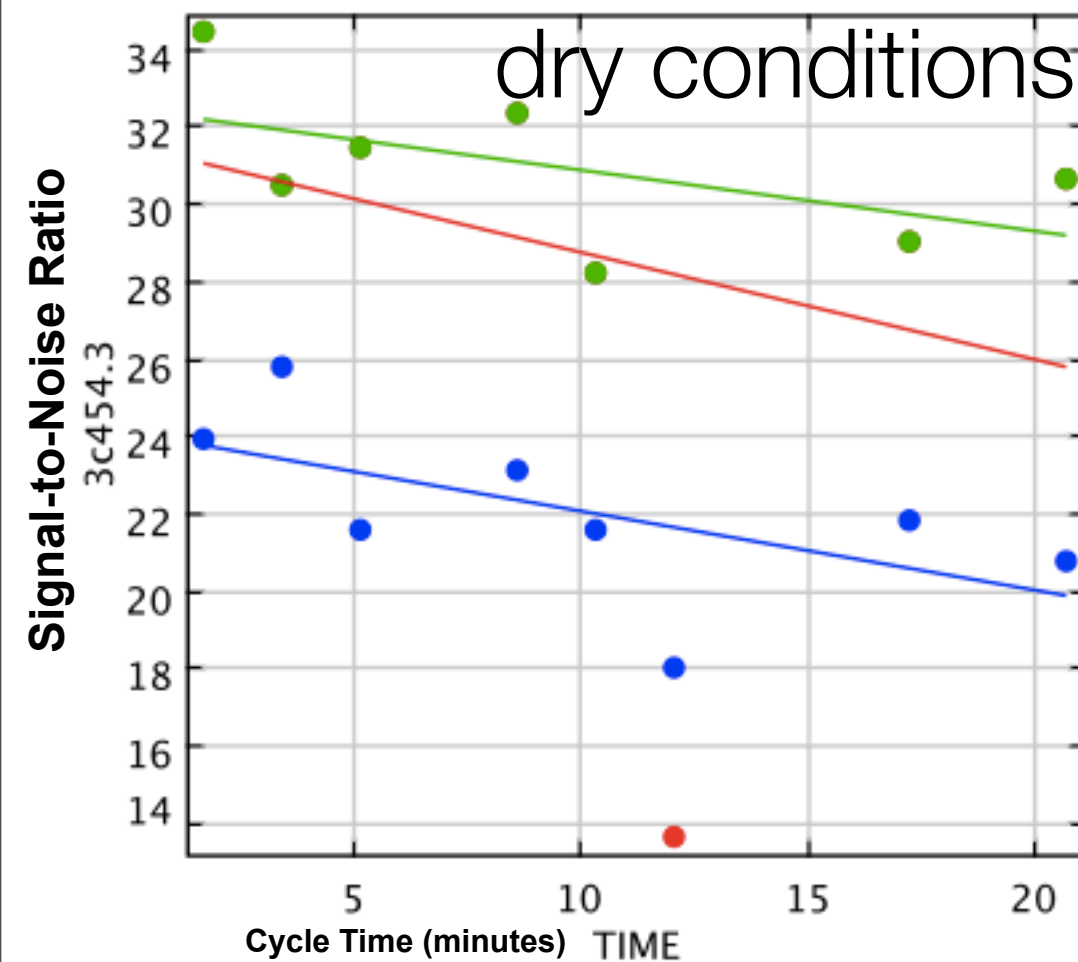
- **Fast switching** (e.g. #262, #403, #404)
 - Track phase fluctuations with a calibrator source.
 - Strong calibrator (optimum at 90 GHz), and closeby to minimize slewing time to target source.
 - r.m.s phase structure function: rms phase increase with the baseline length (up to 16 km!)
 - How to scale the phase from 90 GHz to target freq.?

Fast Switching Test (Master thesis V. Martinez)

-3 QSOs imaging data

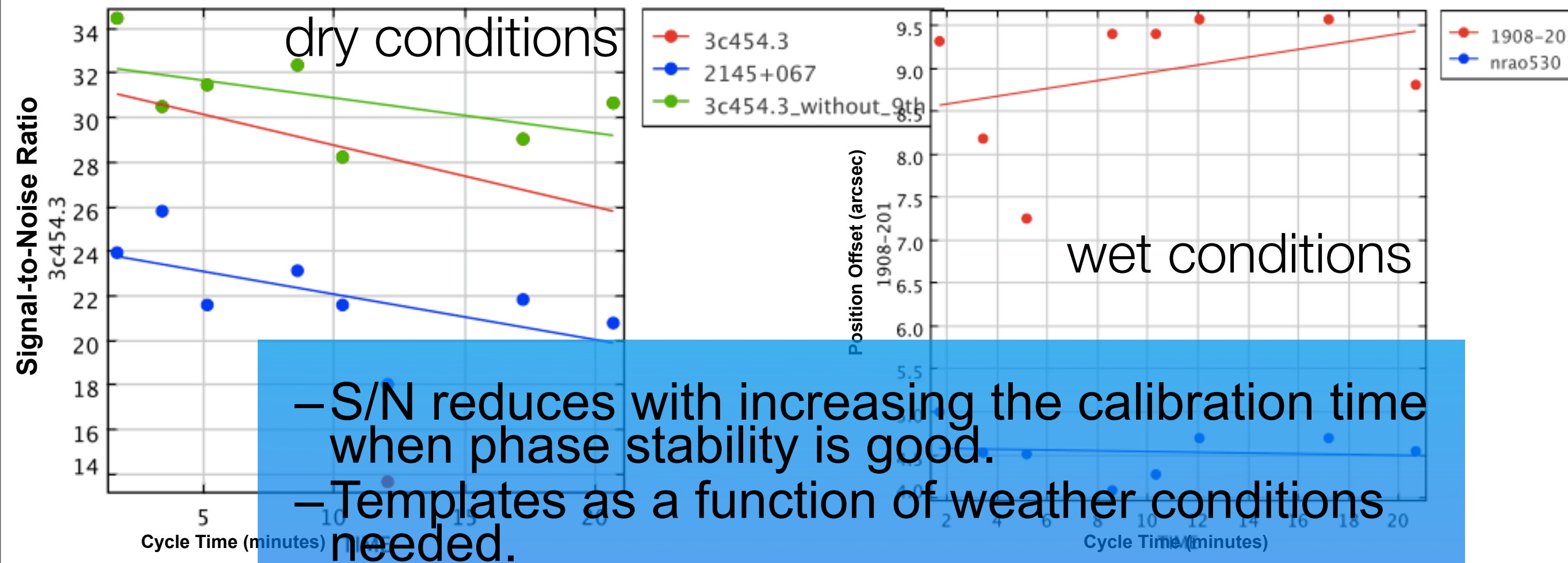
-7 SMA datasets. Maximum baselines ~ 260 m. 345 GHz.

-Cycle= Primary calibrator - QSO1 - QSO 2. 20 sec integration each (5 sec integration time) during 4 hours.



Fast Switching Test (Master thesis V. Martinez)

- 3 QSOs imaging data
- 7 SMA datasets. Maximum baselines ~ 260 m. 345 GHz.
- Cycle= Primary calibrator - QSO1 - QSO 2. 20 sec integration each (5 sec integration time) during 4 hours.



Phase calibration techniques

- Phase Transfer

- Atmosphere:

- Ratio of phases between high freq and low freq:

$$R = (\Delta L_2 / \lambda_2) / (\Delta L_1 / \lambda_1) = f * (v_2 / v_1)$$

- f depends on the AM model. Ex. $f = 0.97$ in Mauna Kea (S. Paine's), $R = 2.97$ for 658.006 / 215.595.

- Non-linearity close to H₂O absorptions (e.g. #404).

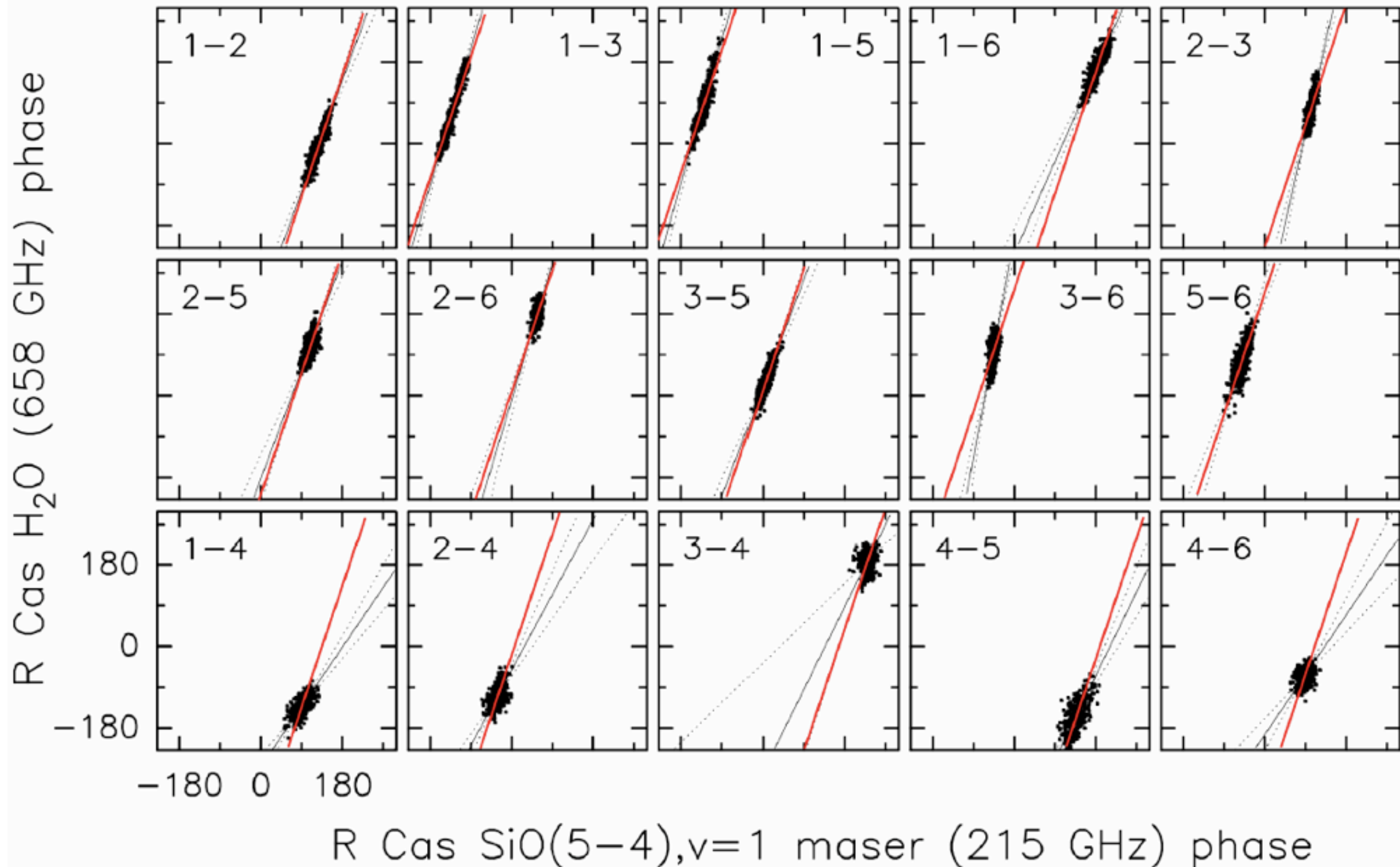
- Instrumental effects: Artificial phase drifts (and jumps)

SMA Phase Transfer test, T. Hunter, December 2005

Dec 2005 @ SMA, T. Hunter

Phase scatter plot over 3.5 hours (15-sec scans)

Expected slope: Red line = 2.97

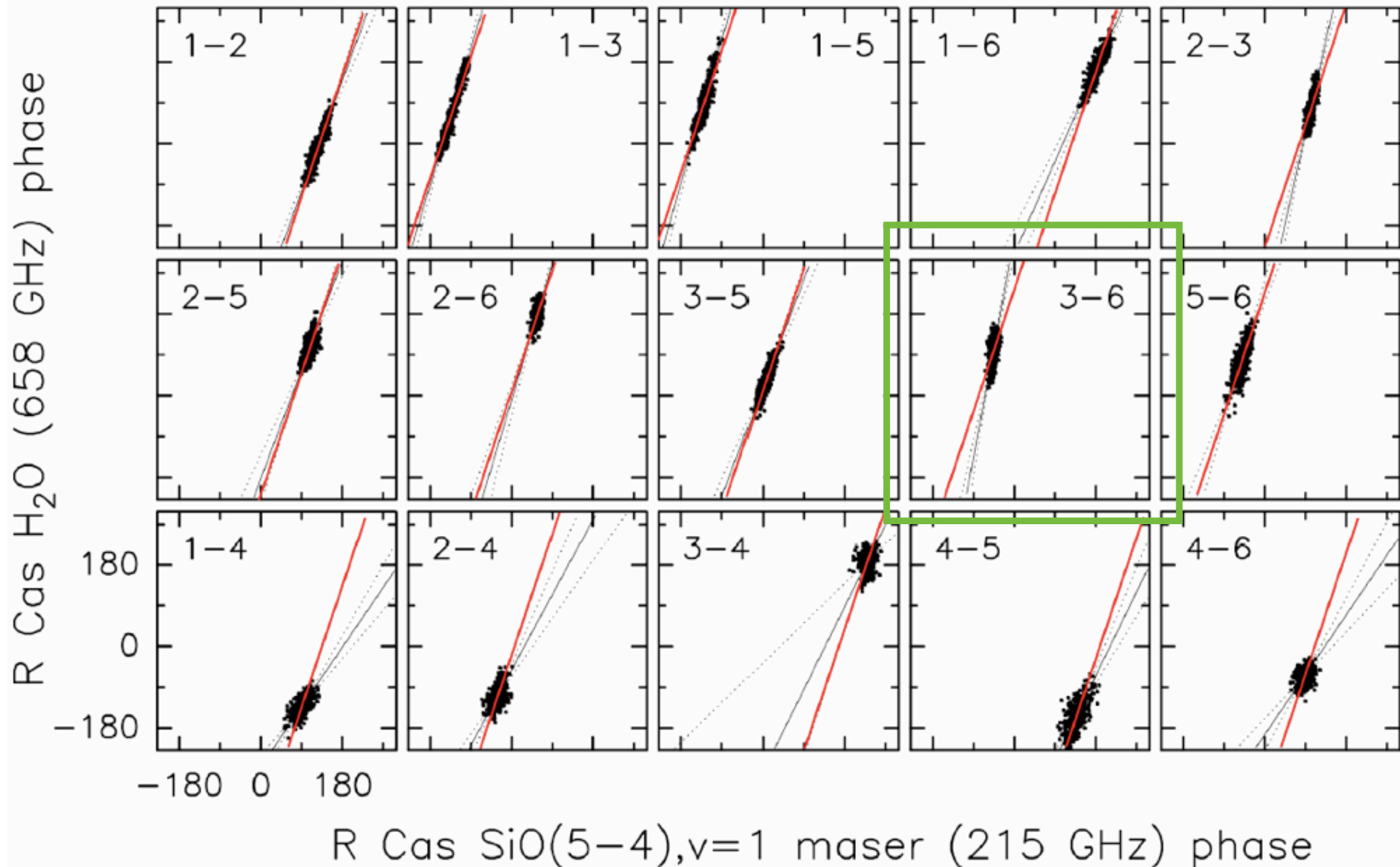


SMA Phase Transfer test, T. Hunter, December 2005

Dec 2005 @ SMA, T. Hunter

Phase scatter plot over 3.5 hours (15-sec scans)

Expected slope: Red line = 2.97

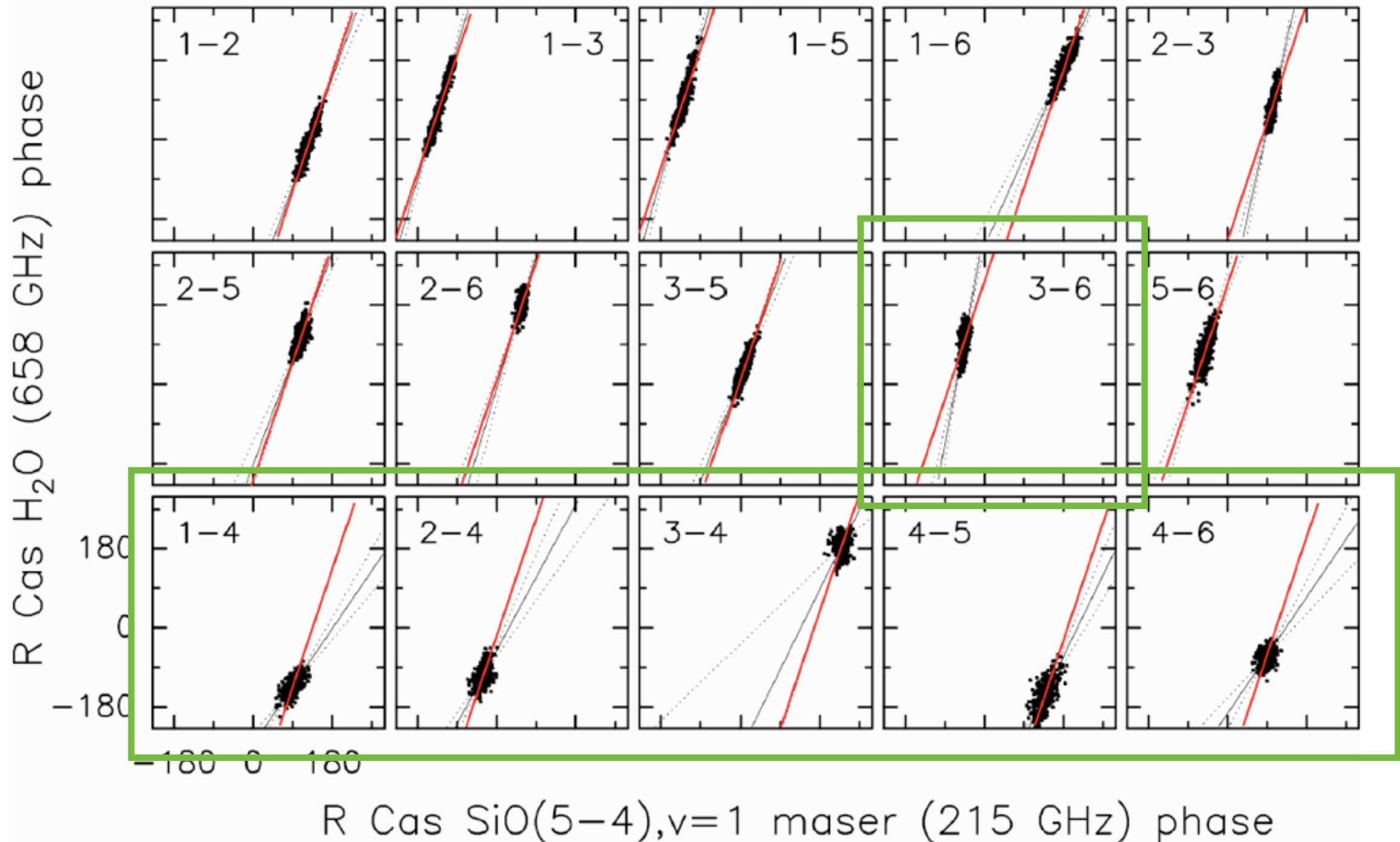


SMA Phase Transfer test, T. Hunter, December 2005

Dec 2005 @ SMA, T. Hunter

Phase scatter plot over 3.5 hours (15-sec scans)

Expected slope: Red line = 2.97



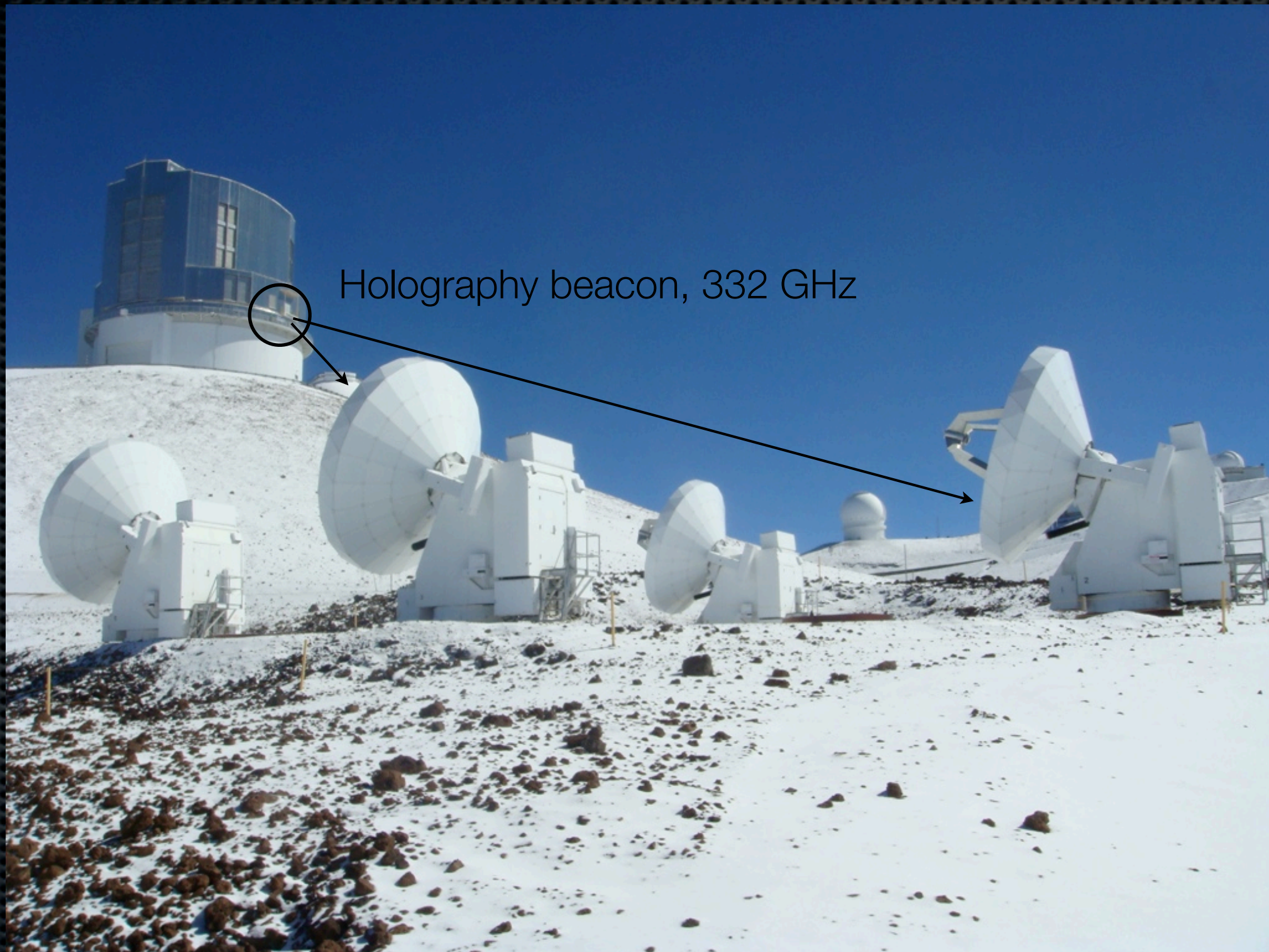
SMA Dual Receiver Beacon Test -- Setup

October 2007



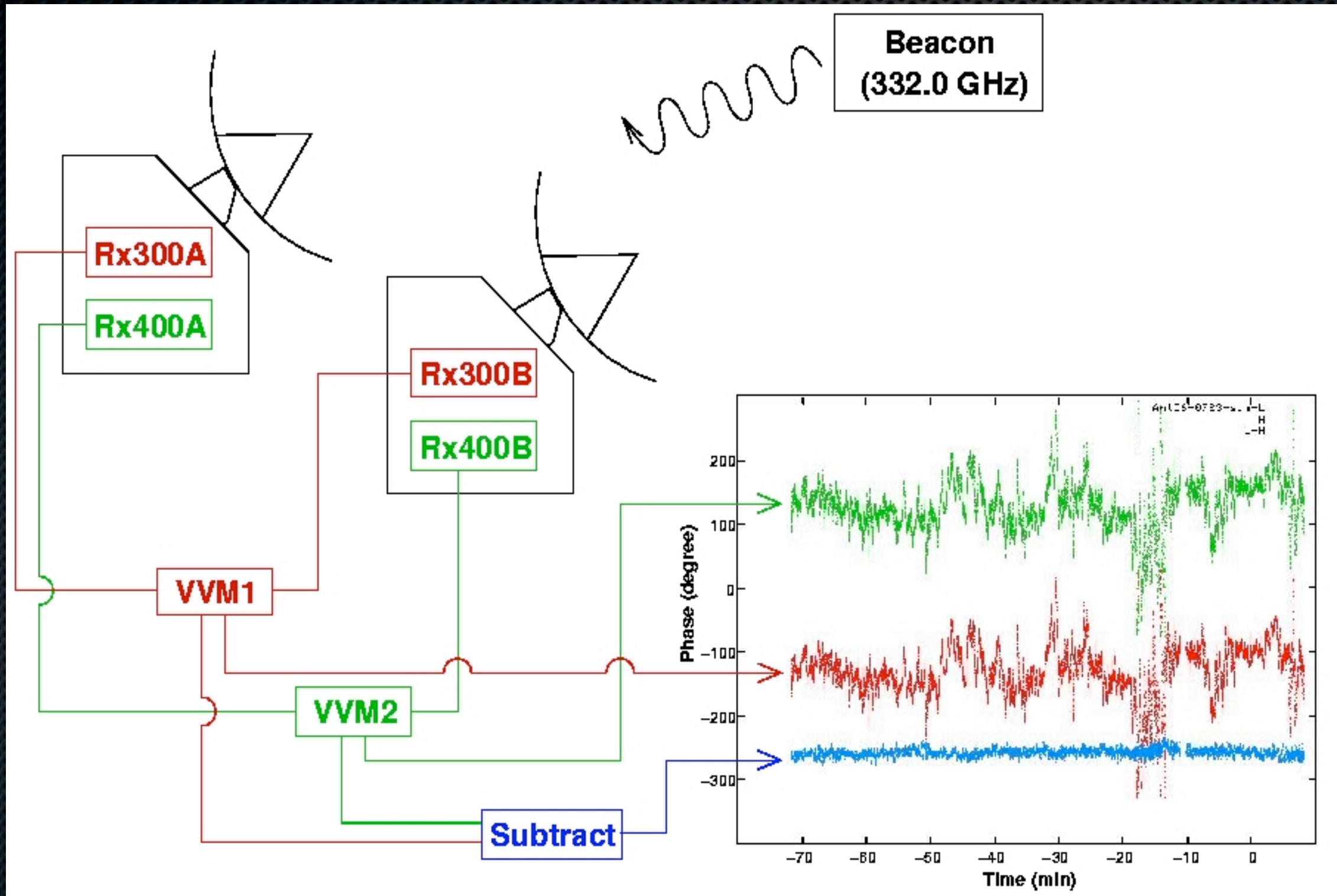
SMA Dual Receiver Beacon Test -- Setup

October 2007

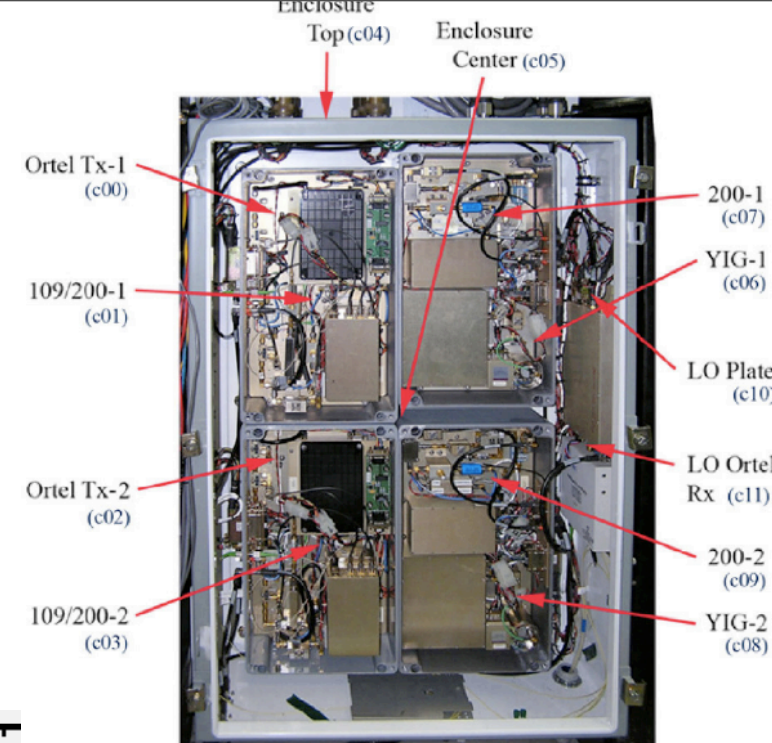


SMA Dual Receiver Beacon Test -- Setup

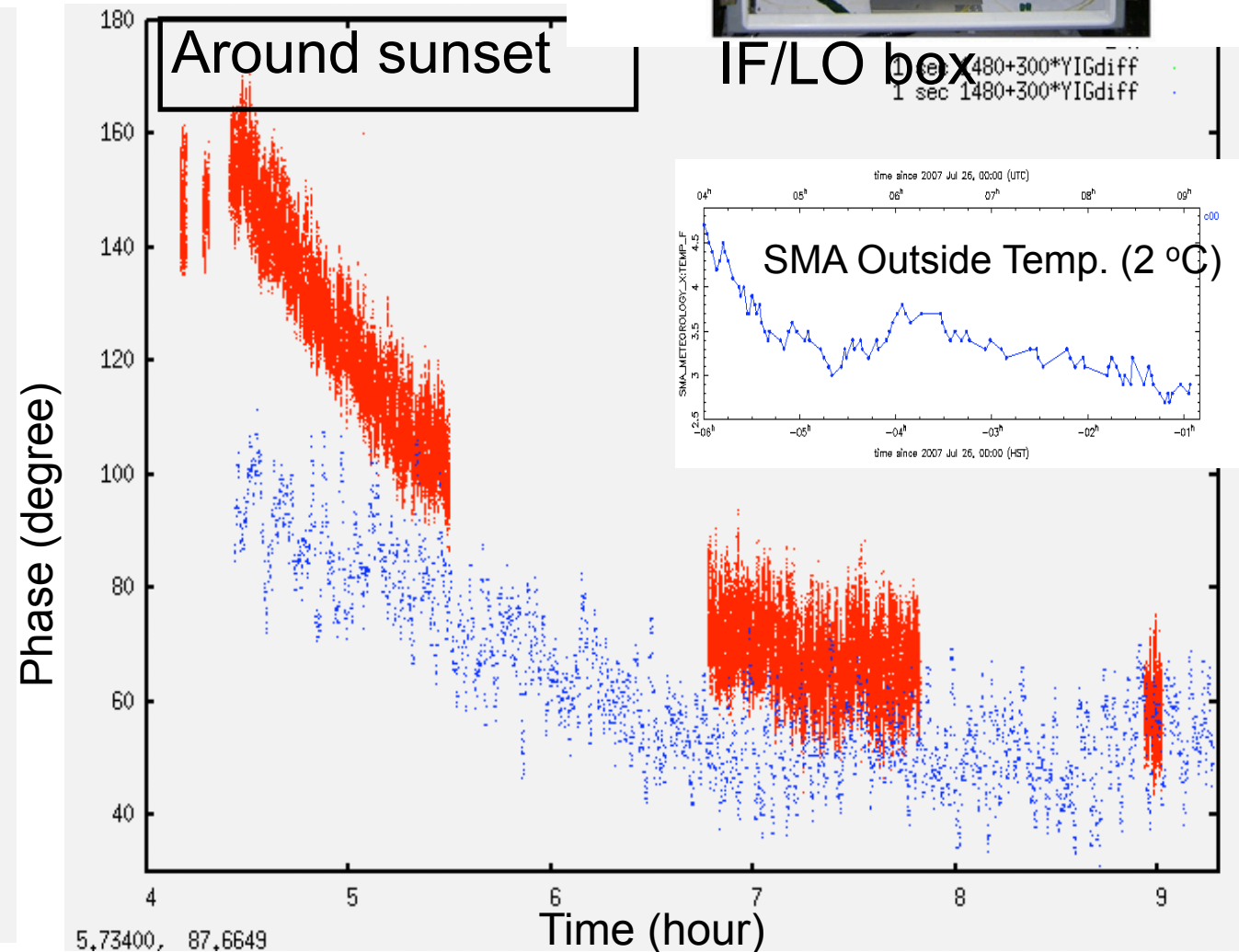
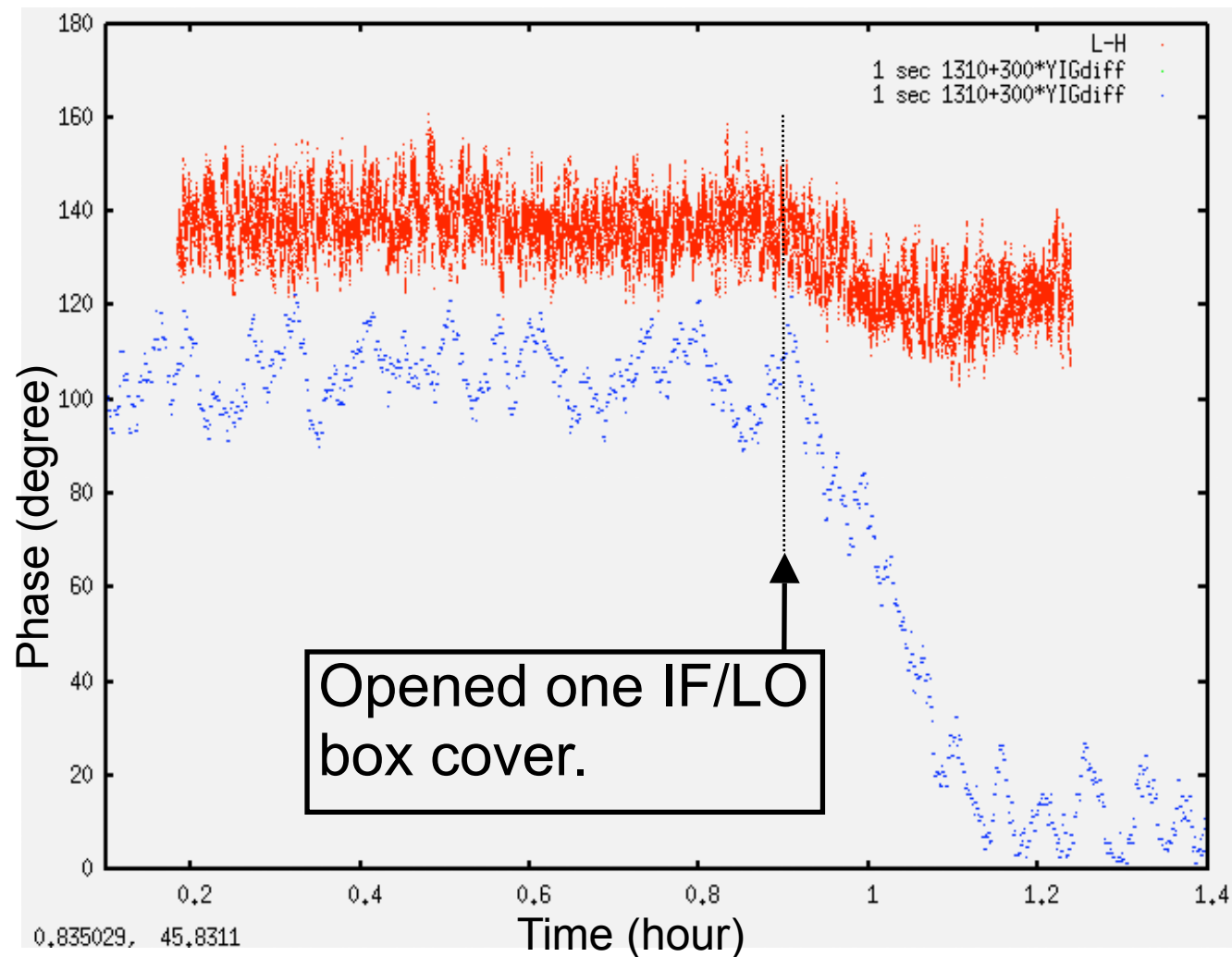
October 2007



Dual Receiver Beacon Test Preliminary Results (1): IF/LO box

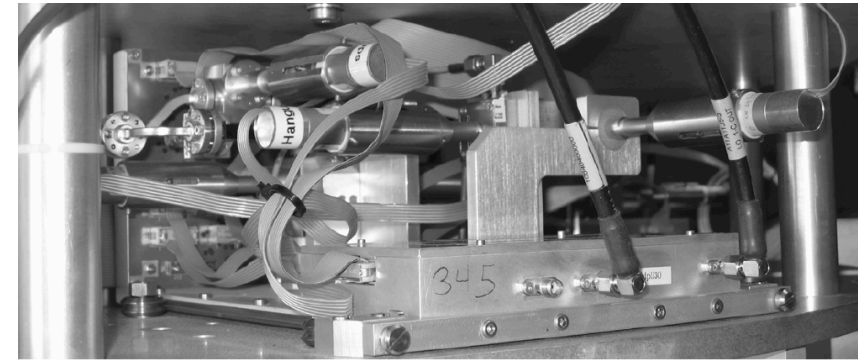


IF/LO box

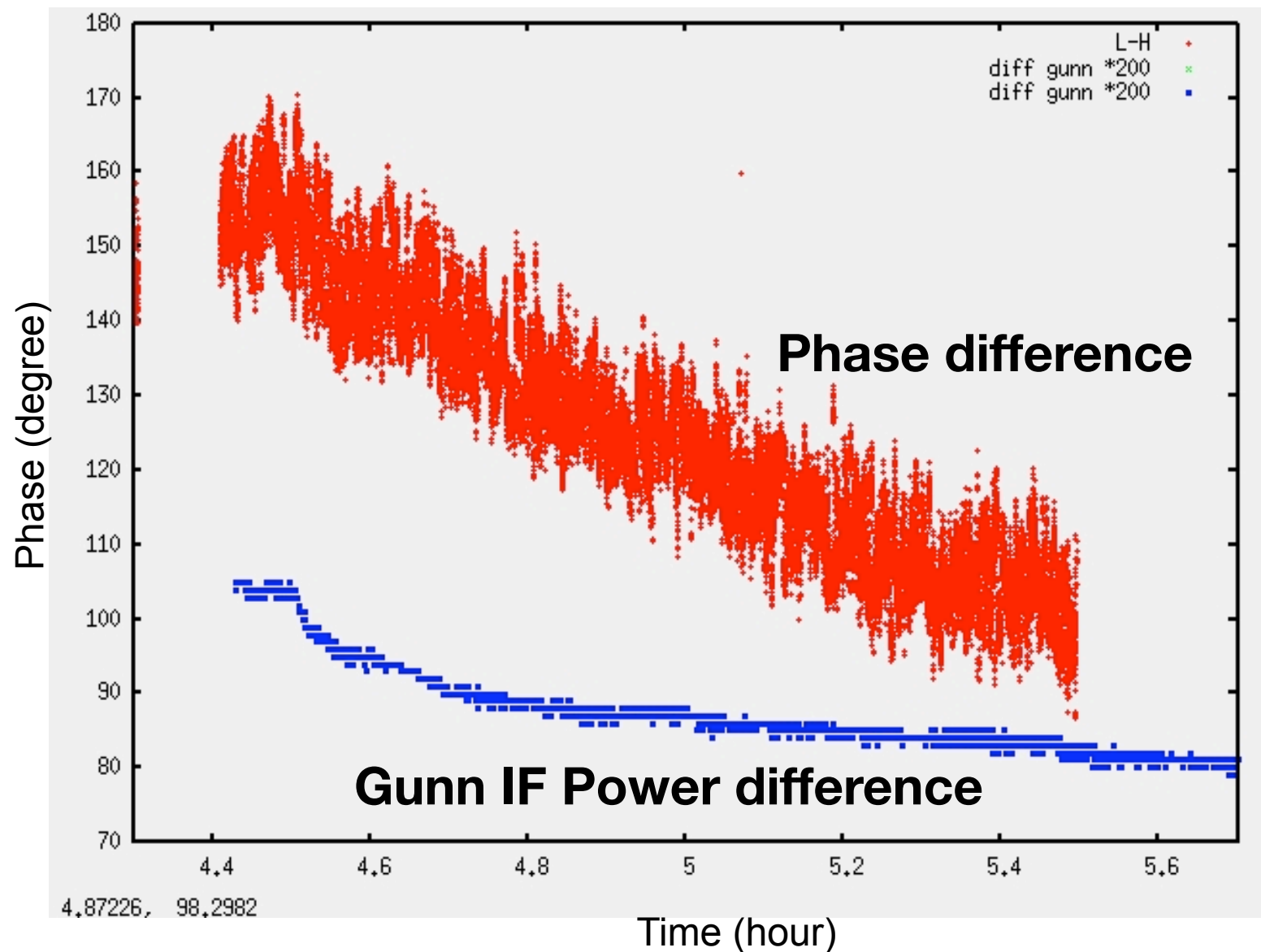


Red data: Phase fluctuation (Rx300A - Rx300B) - (Rx400A - Rx400B)
 Blue data: YIG temperature (Rx300A - Rx300B) - (Rx400A - Rx400B)
 (YIG temp. scale is the vertical scale divided by 300)

Dual Receiver Beacon Test Preliminary Results (2): Gunn LO plate



Gunn LO plate



Red data: Phase (Rx300A - Rx300B) – (Rx400A – Rx400B)

Blue data: Gunn IF power (Rx300A - Rx300B) – (Rx400A – Rx400B)

Actual temperature scale is the vertical scale divide by 200.

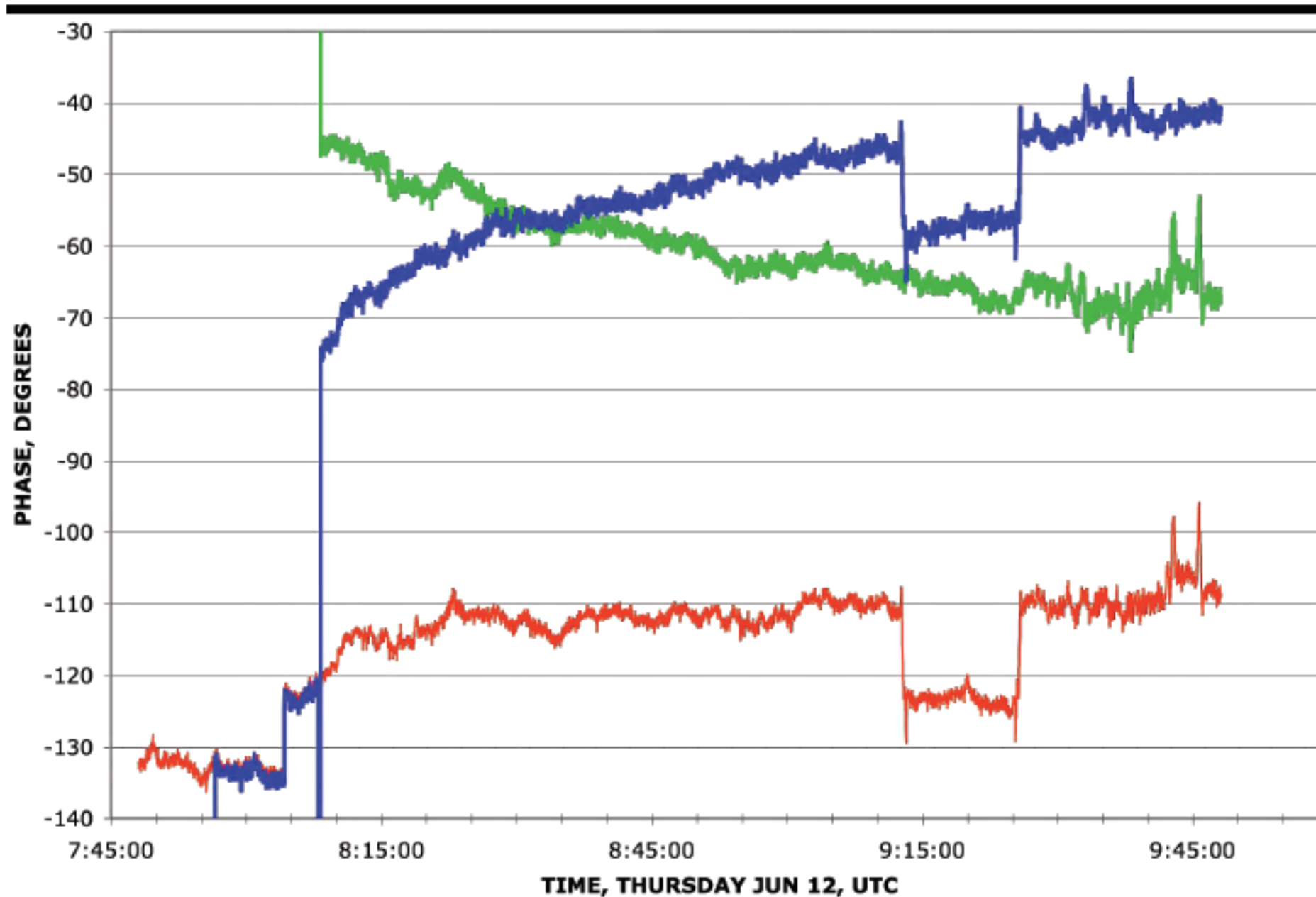
Dual Receiver Test Preliminary Results (3): Azimuth (RA: 0-180°)



SMA Phase Transfer Project



A6-A5 PHASE A1-A5 PHASE A6-A1 PHASE



Dual Receiver Beacon Test

Next Step

- Stabilize temperatures of IF/LO boxes and Gunn plates.
- Install more thermistors in sensible components.
- Study possible correlation of eng. variables and phase drifts and jumps.
- Other effects: azimuth, etc.

Summary

- SMA is suited as pathfinder to study phase calibration of ALMA since covers the mm/submm range and located in a relatively dry place.
- FS: It works in SMA. We are characterizing cycle time with atm. weather and rms structure function.
- PT: It does not work in SMA. We find difficult to correct instrumental effects from the IF/LO box and Gunn plate.
- ALMA calibration scheme is a combination of WVR+FS, but little is planned according to the needed PT, especially instrumental effects. Nice tests for CSV...