ALMA 3D: VO-compliant archive and datacube analysis



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Spanish exploitation of ALMA

Madrid 18 June 2008

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Analysis of the Interstellar Medium of Isolated GAlaxies

OUTLINE

Context: AMIGA project

- Motivation: ALMA, e-Astronomy
- Previous work: VO-archives & tools
- ALMA exploitation: VO-compliant
 - Archive
 - Tools for high level analysis

ASTRONET proposal

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Kapteyn Institute, IAA, Obs. Paris, SVO +

interest from ALMA European Computing Manager

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CONTEXT

AMIGA project:

Analysis of the interstellar Medium of Isolated GAlaxies Starts in 2003 @IAA with funding from PNAYA

Since 2006 Coordinated project (L. Verdes-M) IAA-group + IRAM-30m @ Granada (PI R. Mauersberger)

+ International collaboration:
Obs. Marseille, Obs. Paris, CfA, ASIAA-Taiwan, MPIfA
(Bonn), Univ. Alabama, UMASS, Mc Donald
Observatory, Arcetri, UNAM, IAC, Kapteyn Institute

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CONTEXT

Need for a reference sample of isolated galaxies Multiwavelength statistical study of ISM ~<u>1000 galaxies</u> Build & analyse the catalog (ISM – SF – AGN) Make it public: VO menace with search utilities Results

10 +1 Papers + 1 submitted PhD: 1-2005, 1-2006, 2-2008, 1- 2009, 2- 2010



MOTIVATION: ALMA

Interests:

ISM in galaxies and connection with dynamics, SF, nuclear activity, bars Based on intensive analysis of 3D data: need for powerful soft package for datacube analysis Observational expertise: single dish + interferometry ecm, mm, extension to submm: - intensive use of SMA - calibration tasks for SMA (participation in ALMA CSV)

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MOTIVATION: e-Astronomy

We find:

Few radio data available in archives (not to mention VO...) Optical/IR data more often available, but too diverse queries -VO Essential for multi-λ astronomy VO is part of general context of e-Science: enhanced Astronomy Actions: Start to work on radio-VO: access + exploitation of archives Lead 1st activity to coordinate regional e-Science: e-CA project (e-Ciencia Andaluza) Funded as Proyecto de Excelencia by Junta de Andalucía Involves > 40 research groups + companies

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PREVIOUS WORK: ARCHIVES

DSS-63 Robledo antenna (+LAEFF-INTA)

30m-IRAM

No data model existed for radioastronomical data

RADAMS:

Radio Astronomy DAta Model for Single-dish telescopes
VO-compliant: based on existing IVOA data models + development of new specific standards
Extensible: additional metadata can be provided for different instruments, observing modes, switching modes...

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PREVIOUS WORK: TOOLS

Existing VO-enabled <u>analysis tools</u> are mostly <u>optical</u>
 Existing <u>radio</u> astronomy tools <u>not VO-aware</u>

Need for VO-enabled radio oriented analysis tools

Solution: INTEROP Not producing new soft but adding VO functionalities Intercommunication in the least intrusive way

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PREVIOUS WORK: TOOLS

MOVOIR Development

MOdular Virtual Observatory Interface for Radio-astronomy

Tools: MASSA/MADCUBA (Herschel packages for HIFI, usable with 30m data developed by J. M. Pintado's group)

Data services:

Access to standard FITS imported by the MOVOIR from VO SDSS, HST, MAST, FUSE, IUE, ISO, XMM-Newton, VizieR, AMIGA*, IRAM 30m*, Robledo*...

Applications

Aladin, Topcat, VOPlot, Mirage...

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PREVIOUS WORK: EXPERTISE

Last but not least:since 2005 formation of a team of3 software developers with complementary profiles

- Development of scripts for access to antenna data: filling
 VO data bases
- Design of IVOA standards: data modeling and VO communication protocols (part of Data Model Working Group of IVOA)
- Interfaces for use cases: web services for query and access to VO data

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- Fully integrated in the scientific team
- 1 PhD in 2008 to be followed by a postdoc, 1 PhD starting

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ALMA EXPLOITATION

Key to the success of ALMA: data accessible to the community at large, not only domain of experienced radio astronomers.

This requires access to:

- well documented+intuitive tools to inspect+analyse 3D data
- existing VO tools widely accepted by the community (e.g. Aladin, VOSpec, Topcat, etc)
- complementary data sets at same or different wavelengths



ALMA EXPLOITATION

- ASA Requirements Document states that ALMA Science Archive (ASA) will be VO-compliant.
- Work planned and budgeted: core functionalities
- Support for proposals preparation via the ARCs
- Automatic processing: fully calibrated science products
- Data reduction: CASA
 - Transformation UV-image
 - Spectral line analysis, fitting, catalogs
 - Fundamental image structure: cubes
 - Processing of cubes (integration, rotation, filtering, clipping) Visualization + specific tools for SS objects and pulsars

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ALMA EXPLOITATION: ARCHIVE

Planned collaboration with ESO-Archive team (pending on finantial support for 1 FTE):

- support from head of the Archive
- inputs from Archive Users
- interaction with IVOA Data Modeling Working Group

-Characterization of the ALMA Archive:

- study the relationship between Archive XML Schema, SQL structure, and ALMA Science Data Model (ASDM)

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- mapping between the ASDM and VO Data Models

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ALMA EXPLOITATION: ARCHIVE

-Development of a Radio Data Cube Data Model (RDCDM) suitable for the ASA, to be submitted for approval and discussion to the DMWG

Development of a suitable IVOA data model for radioastronomical data cubes

-VO services:

- analysis of ASA Requirements draft Use Cases stating which use cases can be provided by already existing VO services

- VO spectral and image services will be deployed and tested

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High-level analysis tools for 3D data

- ALMA not expected to have them (ALMA community day 2007, 3D-2008 meeting last week)

- GIPSY (Groningen Image Processing System, developed at Kapteyn AI) one of oldest + most powerful systems available

GIPSY upgrade and integration in the VO, full compatibility with ALMA data, usability in order to make it available to a larger user base Collaboration IAA, Kapteyn Institute, SVO and Obs. Paris

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Fig. 6. a) Contour map of the HI column density distribution in NGC 3642 overlapped on the *R* image. The contours are 1.1, 3.4, 5.6, 7.9, 10.2, 12.4, 14.7, 16.9, 19.2 and 21.4×10^{20} atoms cm⁻². b) Greyscale map of the HI column density distribution with contours as in a). The main features are marked as dots. The synthesized beam $(21.4 \times 18.4 - \alpha \times \delta)$ is plotted in the upper left of both panels.

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Fig. 7. a) Map of the first-order moment of the radial velocity field where both iso-velocity contours and greyscale are shown for clarity. The scale goes as in the wedge, where the numbers indicate heliocentric velocities in km s⁻¹. The contours go from 1532 to 1602 km s⁻¹ with a step of 5 km s⁻¹. The main spiral features are marked as dots. The straight line indicates the direction of the position-velocity cut shown in Fig. 8. The beam size is $21.^{''}4 \times 18.^{''}4$ and is plotted in the upper left. b) Central part of the velocity field shown in a). The major and minor axis directions are indicated, and a cross indicates the optical center position (see Sect. 3.2).

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Obtained with ROTCUR task



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Obtained with GALMOD task

Fig. 10. First-order moment of the modeled channel maps obtained with the geometrical parameters plotted in Fig. 9.

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Definition of functional and technical specifications:

Use cases will be elaborated by scientists via a survey

Intuitive graphical user interfaces: Needed to fully exploit the potential of GIPSY

Ensure data compatibility, in particular with ALMA archive: Integration in the data environment of ALMA and the CASA core system Implementation of VO tools:

interoperability ensures communication with any data provider or service included in a VO-Registry (catalogs, images, spectral libraries, etc)

Coordinated integration of GIPSY + ALMA archive in VO

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CONCLUSIONS

Aimed contribution to

ALMA VO-compliant archive

First high-level, friendly, VO-aware analysis package for radio
 3D data, applicability to multiλ datasets fully compatible with
 ALMA

IVOA standards to cube data models

Non "measurable" results: Formation of a group where developers of radio-VO software (few in the world!) for ALMA work in **direct contact**

extragalactic scientists preparing to make the best use of ALMA

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